

CHIME

PROBING THE RADIO TRANSIENT UNIVERSE

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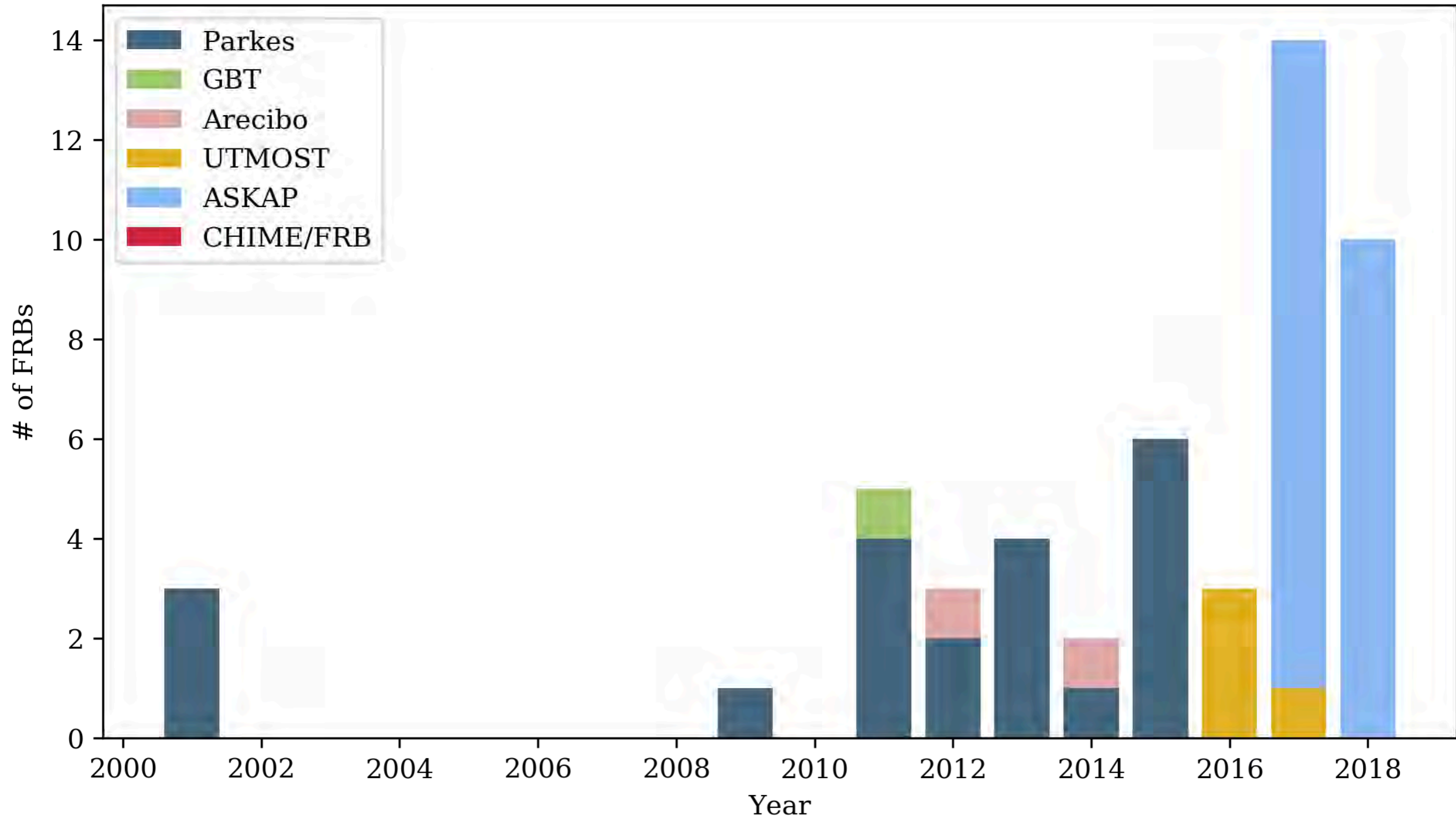
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CHIME PARAMETERS

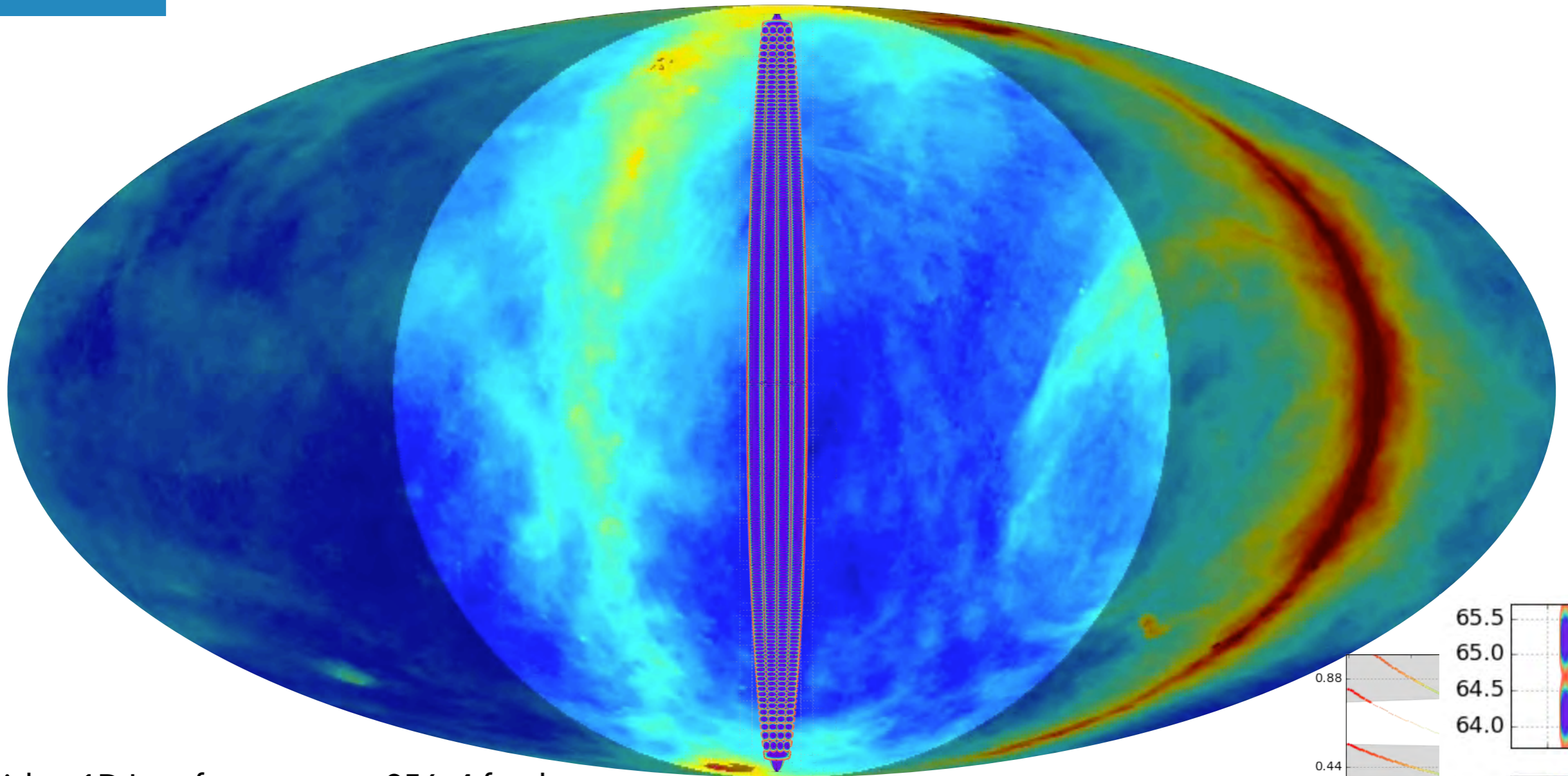
- ▶ 4 Cylinders - 20m x 100m each
- ▶ 1024 Dual Polarization Feeds

Bandpass	400 MHz	800 MHz
21 cm Redshift	2.5	0.8
Beam Size	0.52°	0.26°
E-W FoV	2.5°	1.3°
N-S FoV	~100°	
λ	0.75m	37.5cm

chime-experiment.ca



L0



1 Dish + 1D Interferometry on 256x4 feeds

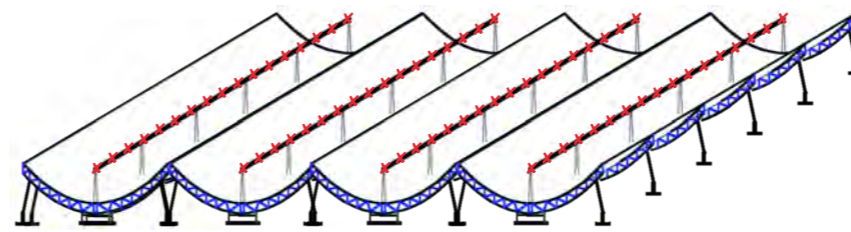
FFT Beam-forming (Tegmark & Zaldarriaga, 2008, 2010)

Hybrid Beam-forming (Cherry Ng et al 2017, 2018)

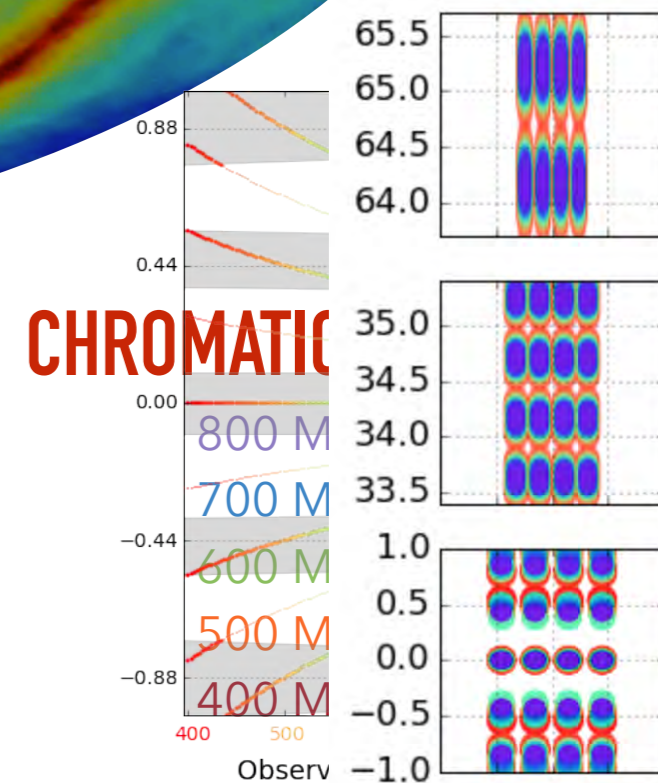
Primary Beam for Single Feed

Hybrid Beam-forming formed beams [E-W]

Sky Coverage ~250 sq. Degrees



@Cherry Ng



Realtime Backends



CHIME/Cosmology

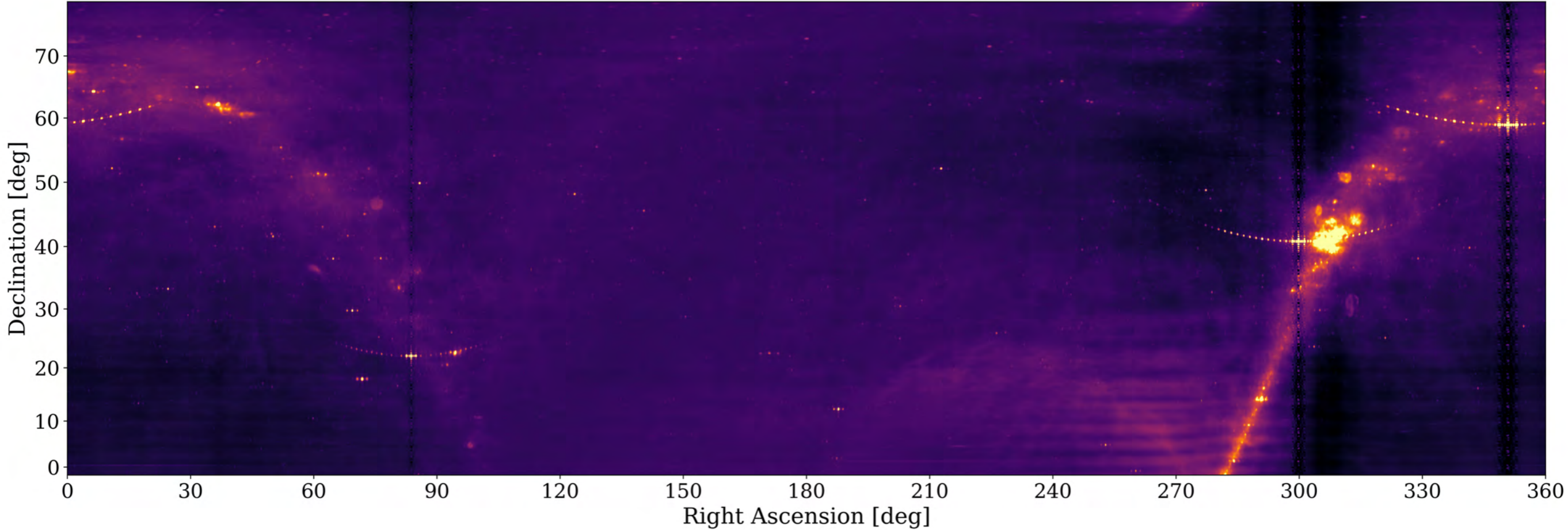
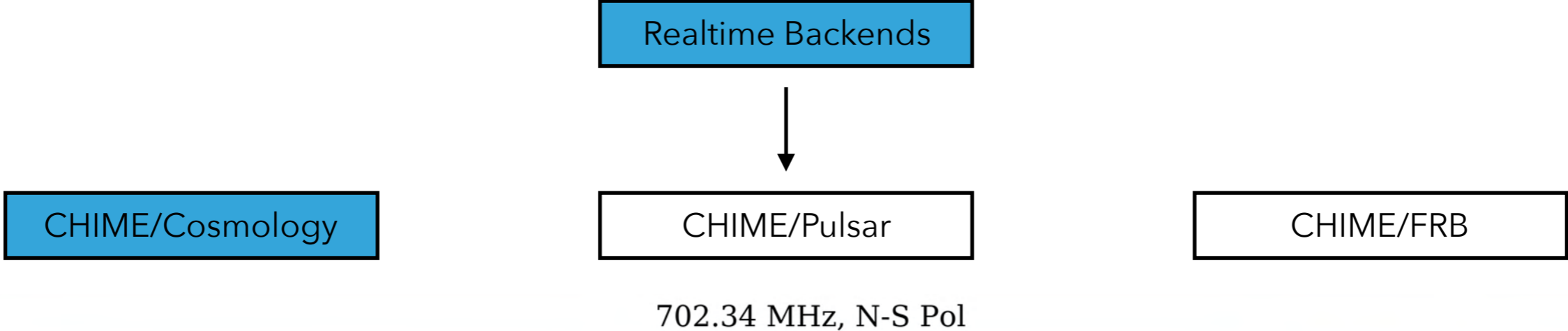
- ▶ Full N^2 visibility matrix
- ▶ 10s cadence
- ▶ 210 TB/day
- ▶ Real time flagging & gain calibration
- ▶ Data compression through redundant baselines (1 TB/day)

CHIME/Pulsar

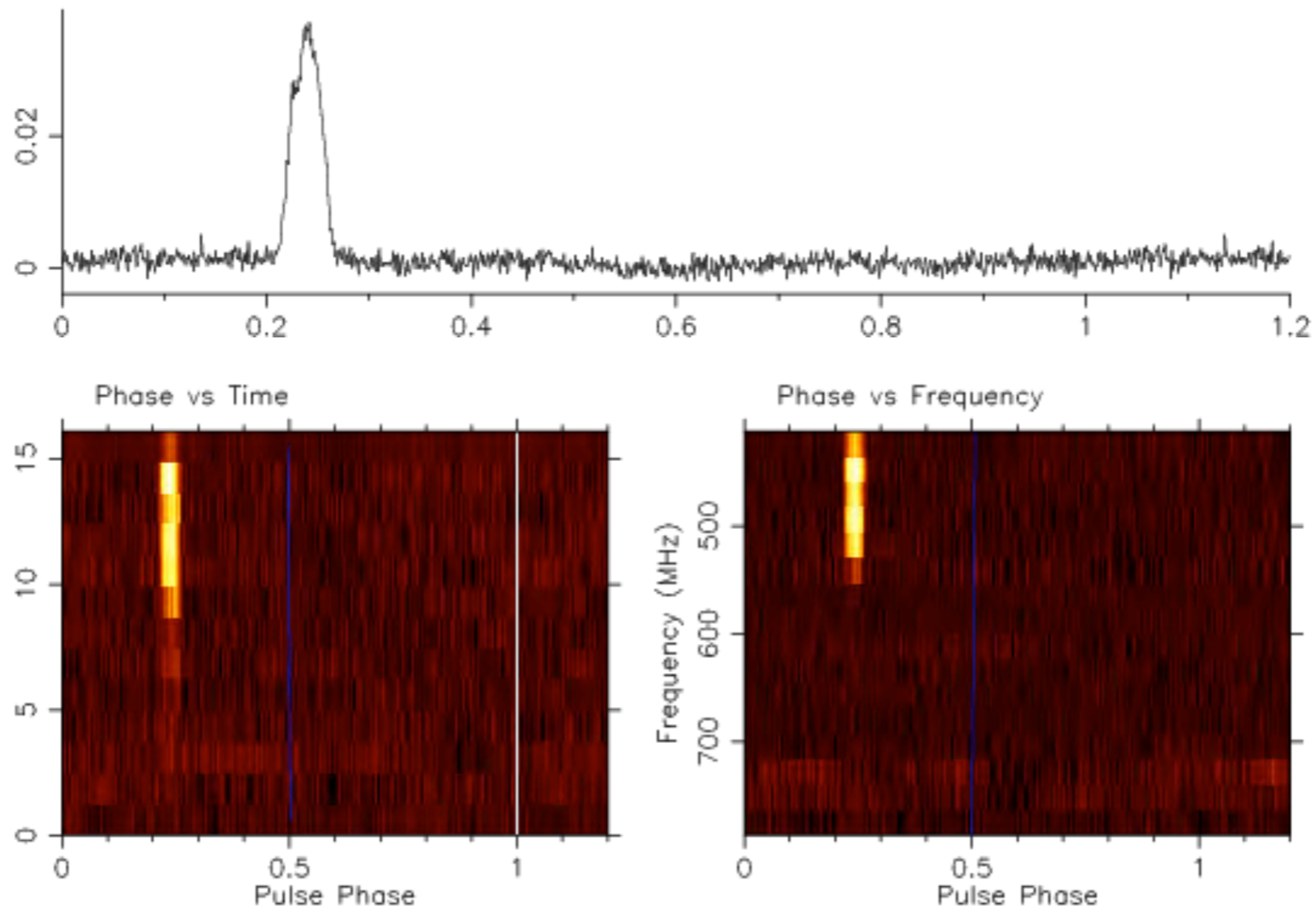
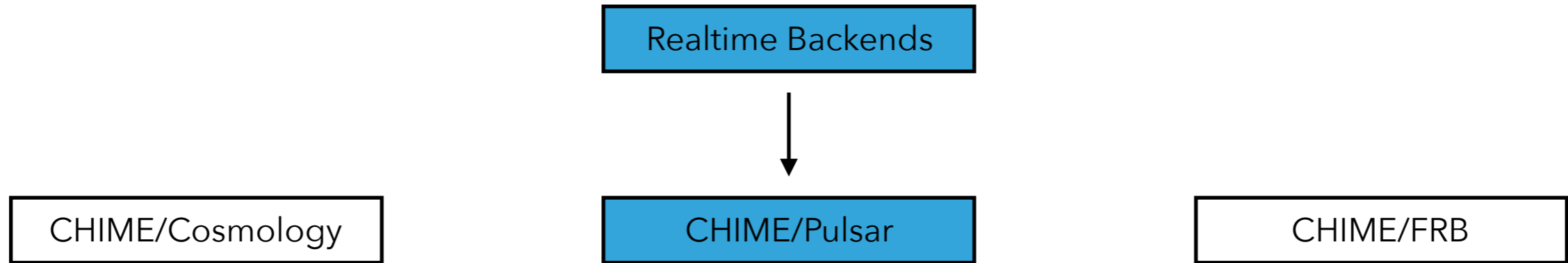
- ▶ 10 tracking beams
- ▶ Track galactic pulsars daily
- ▶ 2.56 μ s cadence, 1024 freq. channels, 2 polarizations
- ▶ 10 x 6.4 Gbps
- ▶ Data compression through folding – 672 GB/day

CHIME/FRB

- ▶ 1024 FFT-formed stationary intensity beams
- ▶ 0.983 ms cadence
- ▶ 16384x 24.4 kHz channels
- ▶ 130 Gb/s (intensity)
800 GB/s (baseband)
- ▶ 1.5 PB/day

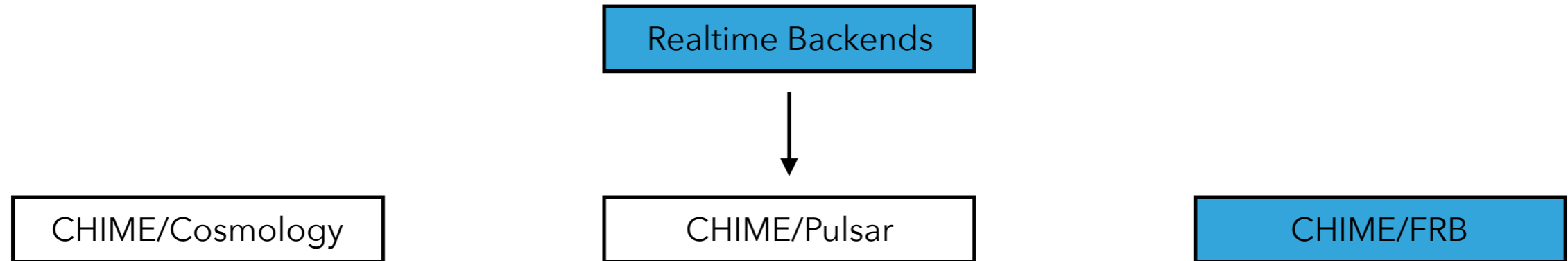


CHIME Cosmology Dirty Map



See Arun Naidu's talk 11:55 am tomorrow

Cycle through all the Northern Hemisphere pulsars in ~10 days!

**L0**

FFT Beamforming + Upchannelization

L1

RFI Excision + Incoherent Dedispersion

L2-L3

MultiBeam Analysis + Science Actions

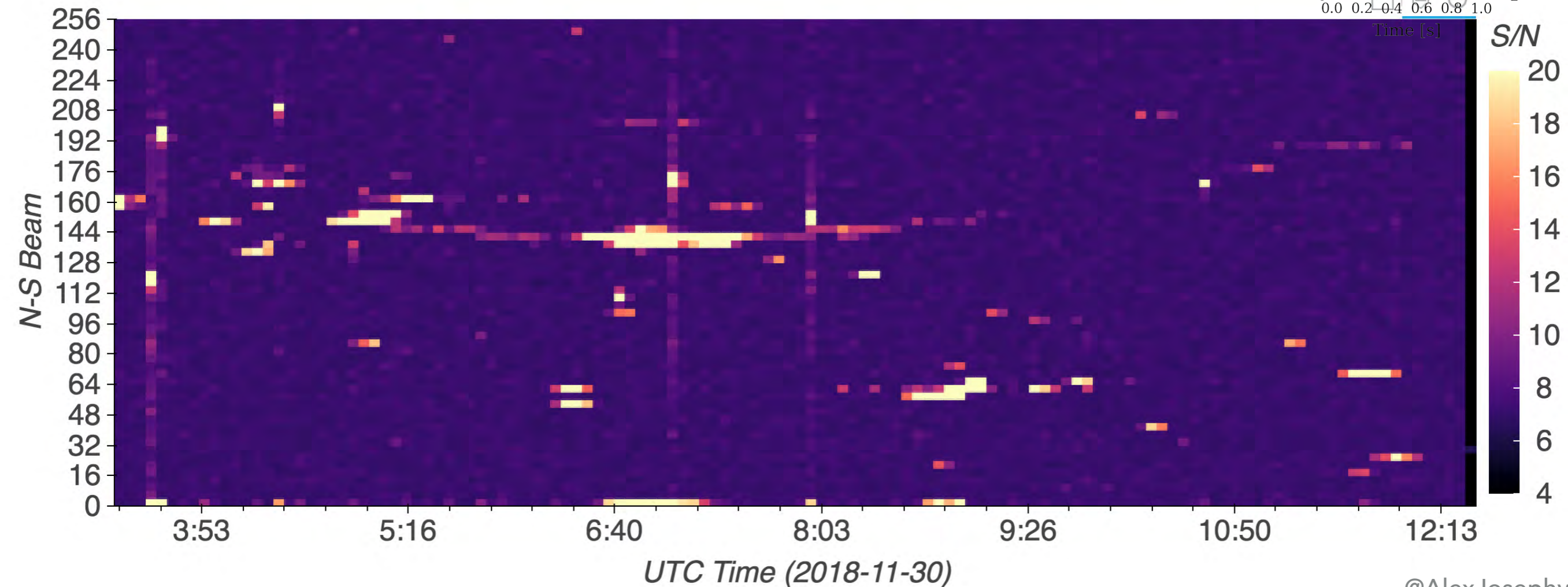
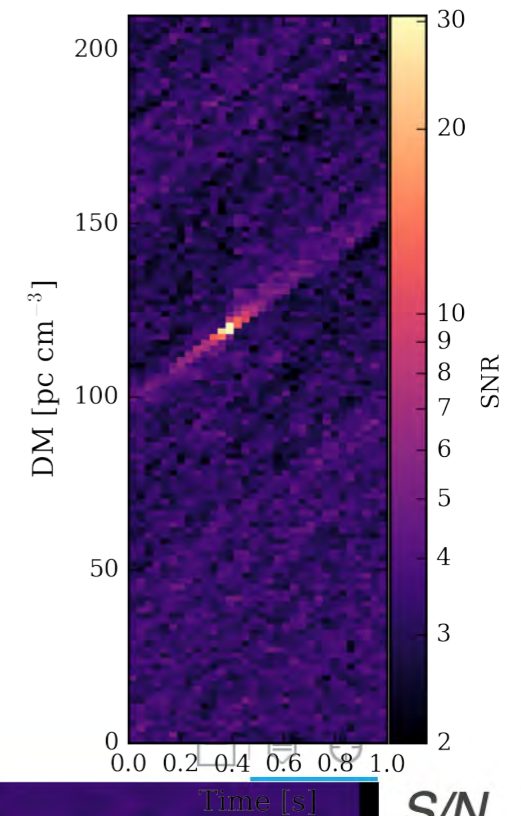
L4

Databases + Offline Processing

L2-L3

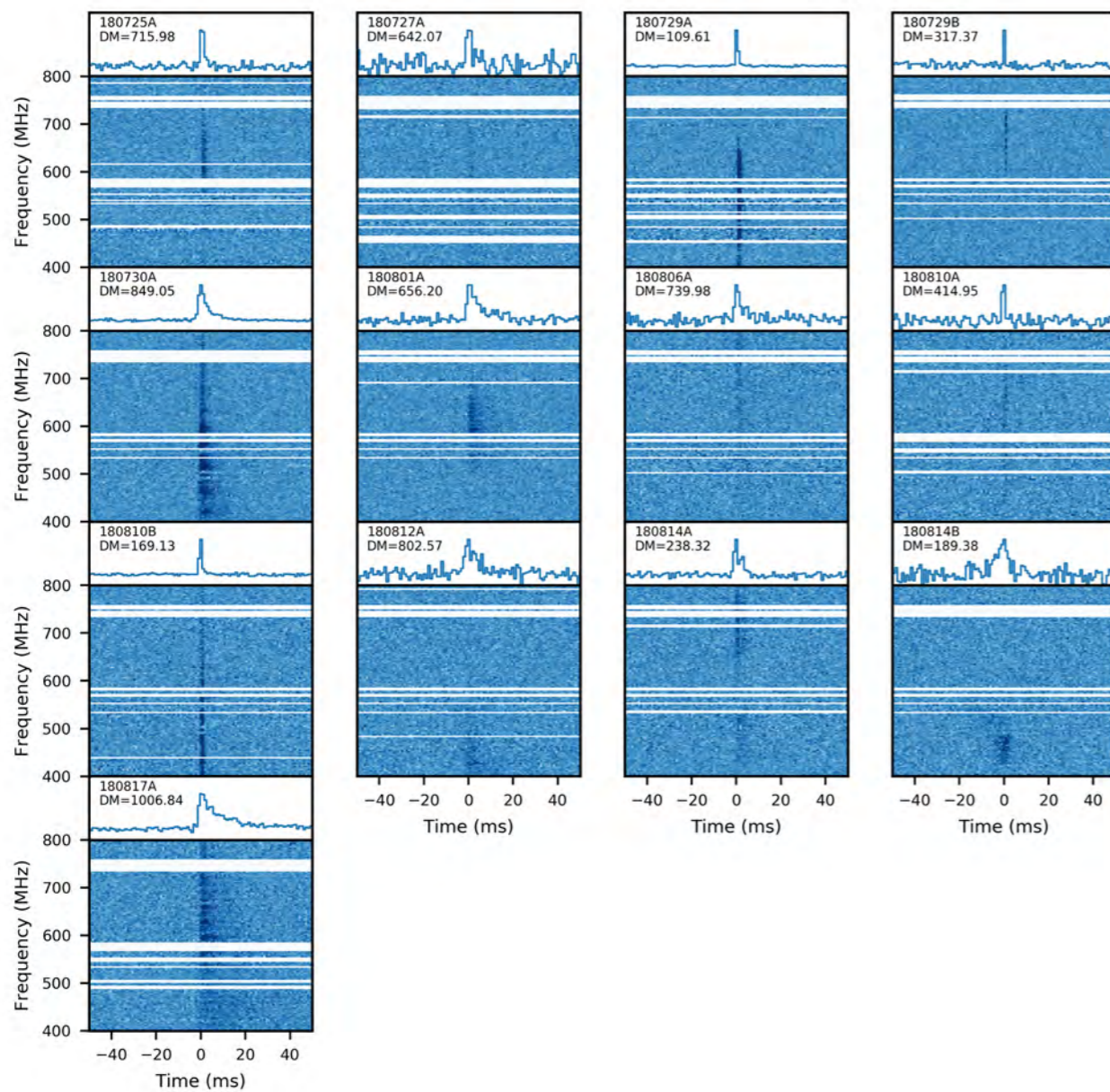
SCIENCE PIPELINE

- ▶ Manage 268.43 Billion Triggers/s
- ▶ Find the one trigger which is the FRB!



SCIENCE RESULTS

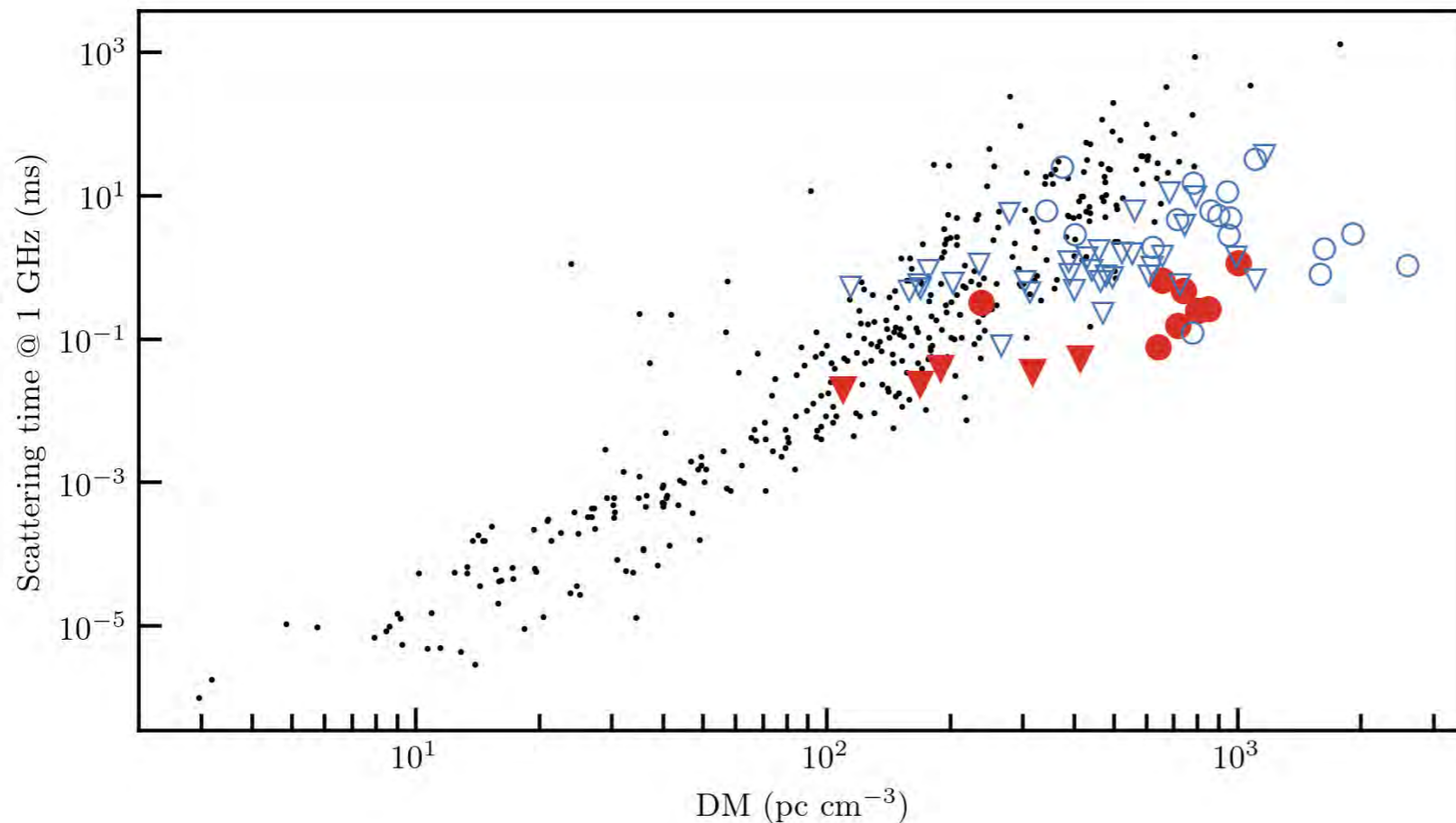
13 NEW FRBS



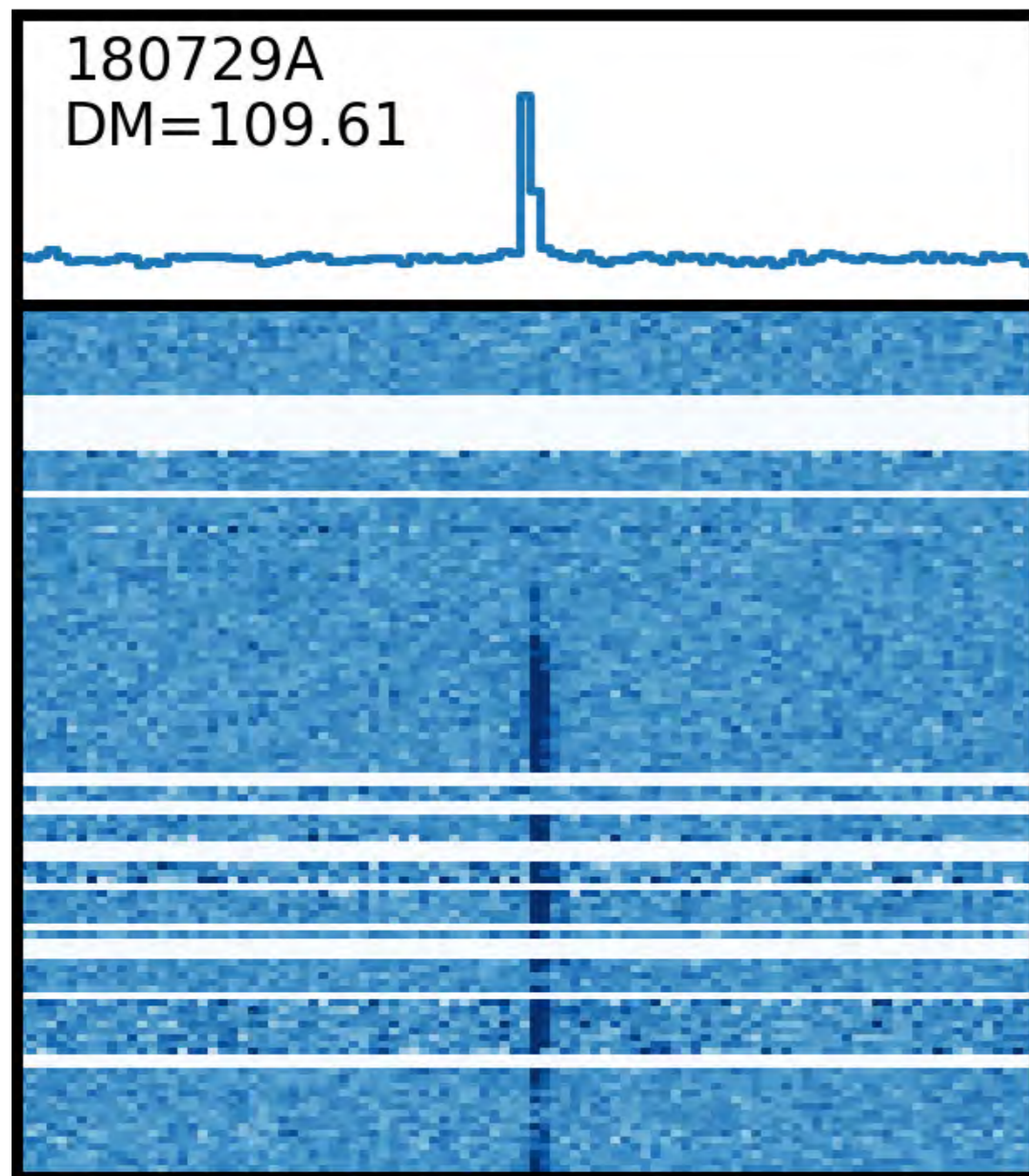
FRB SCATTERING TIMESCALES

7/13 have $t_{\text{scatt}} > 1$ ms

Scattering hard to explain with just spiral arms \rightarrow Needs extra scattering sources

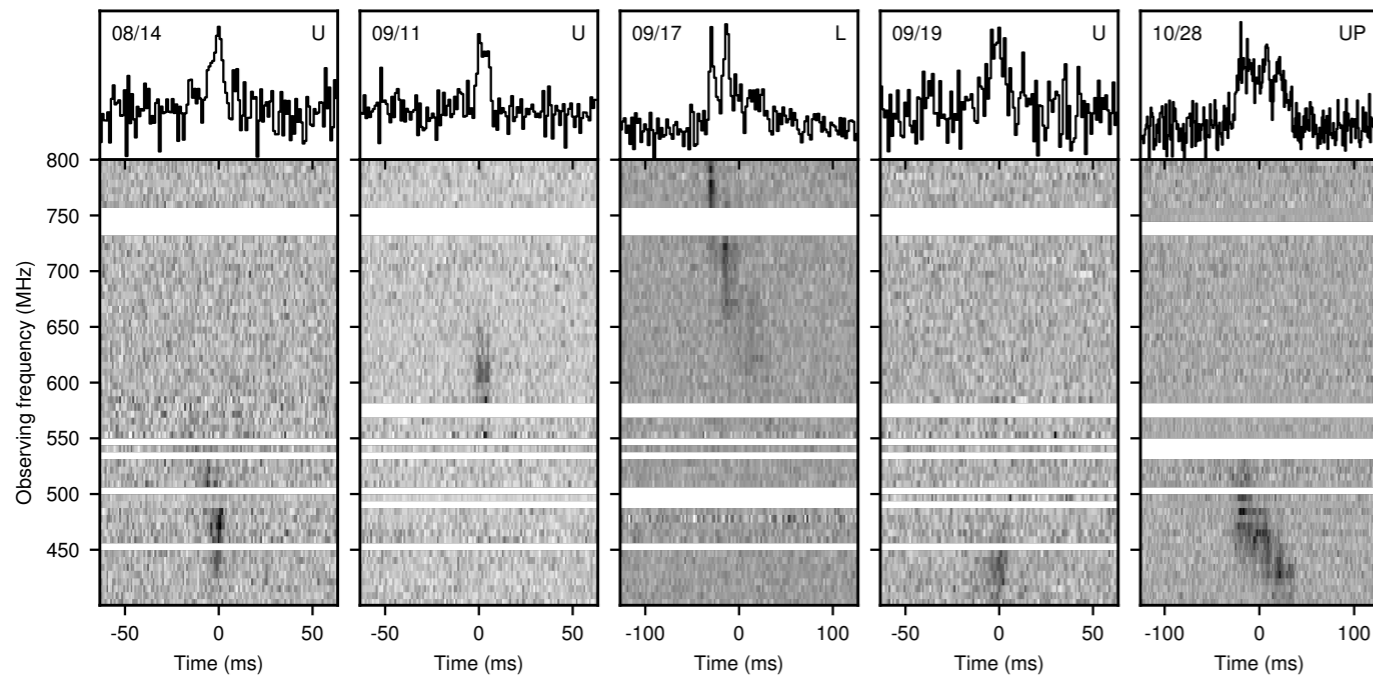


LOWEST DM FRB



- ▶ $DM = 109 \text{ pc/cm}^{-3}$
- ▶ MW: $23\text{-}30 \text{ pc/cm}^{-3}$
- ▶ No scattering
- ▶ $Z < 0.1$, $d < 440 \text{ Mpc}$
- ▶ Assume Host $DM \sim 45 \text{ pc/cm}^{-3}$, $z \sim 0.017$, $d \sim 75 \text{ Mpc}$
- ▶ Many plausible galaxy counterparts, some with radio sources, no bright galaxy

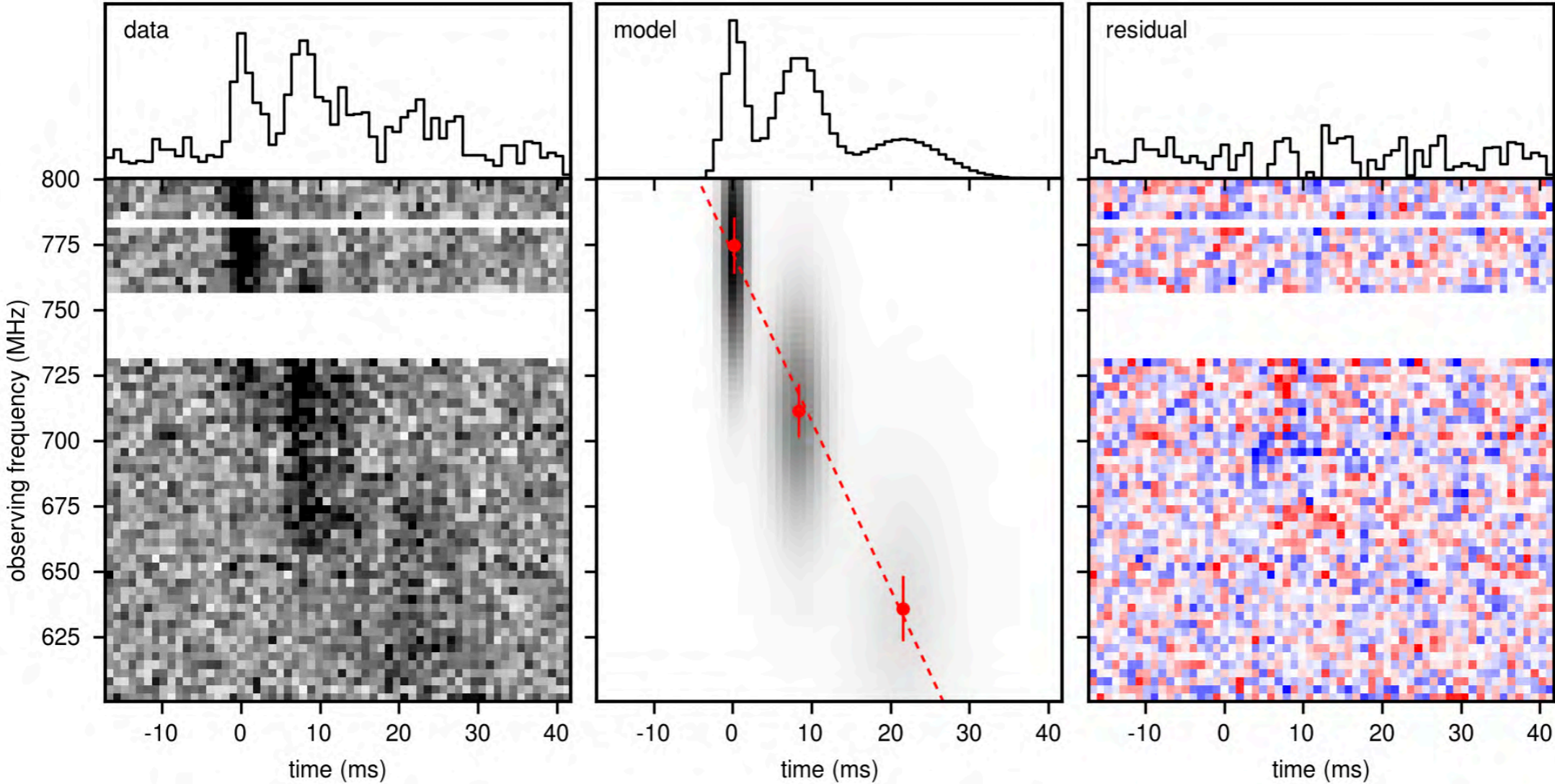
CHIME REPEATER (R2)



RA	DEC	DM pc cm ⁻³	Fluence Jy ms	DM(Gal) pc cm ⁻³
63°4h22m	+73°	189.4	3-60	~80 - 100



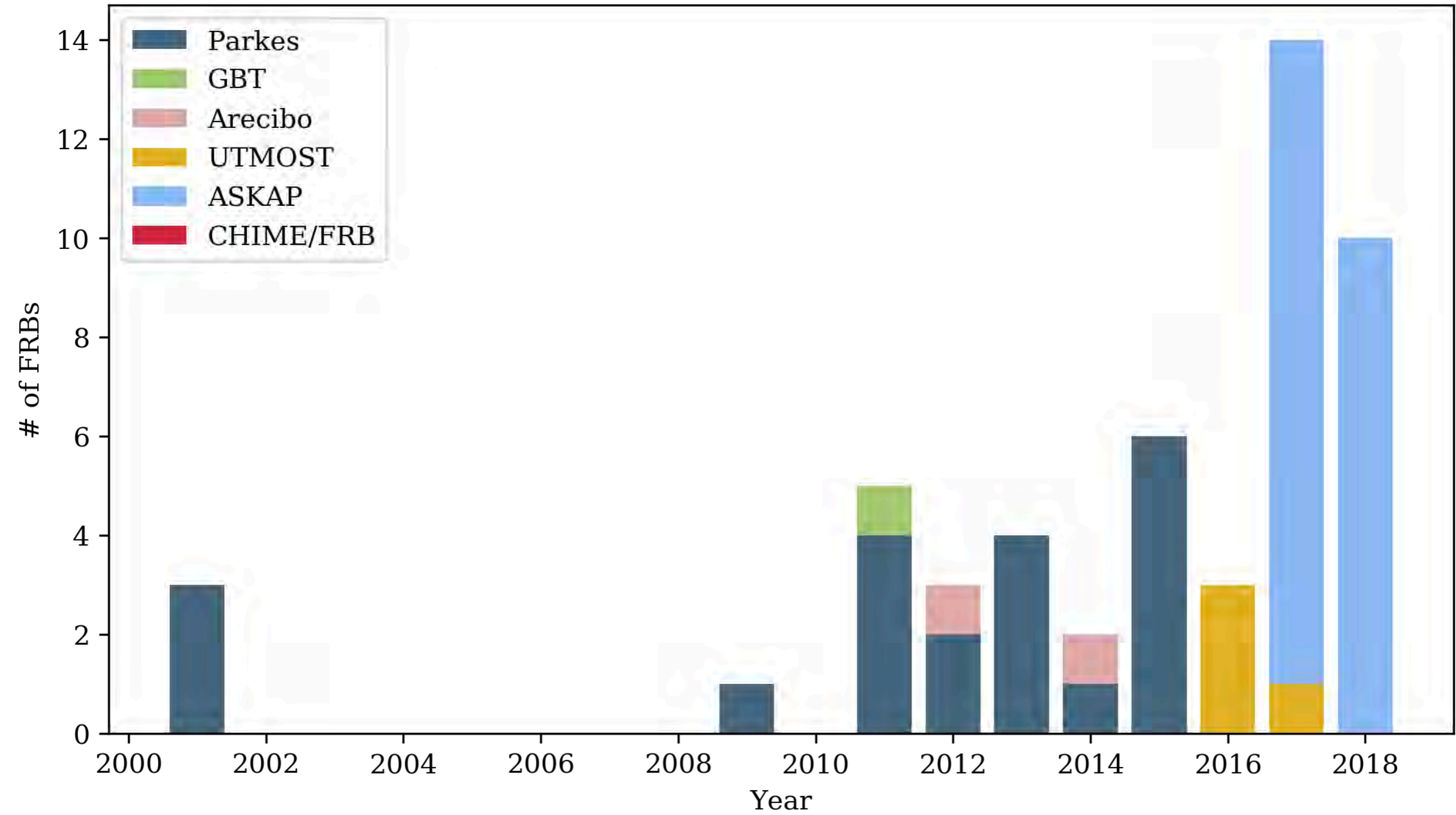
CHIME REPEATER

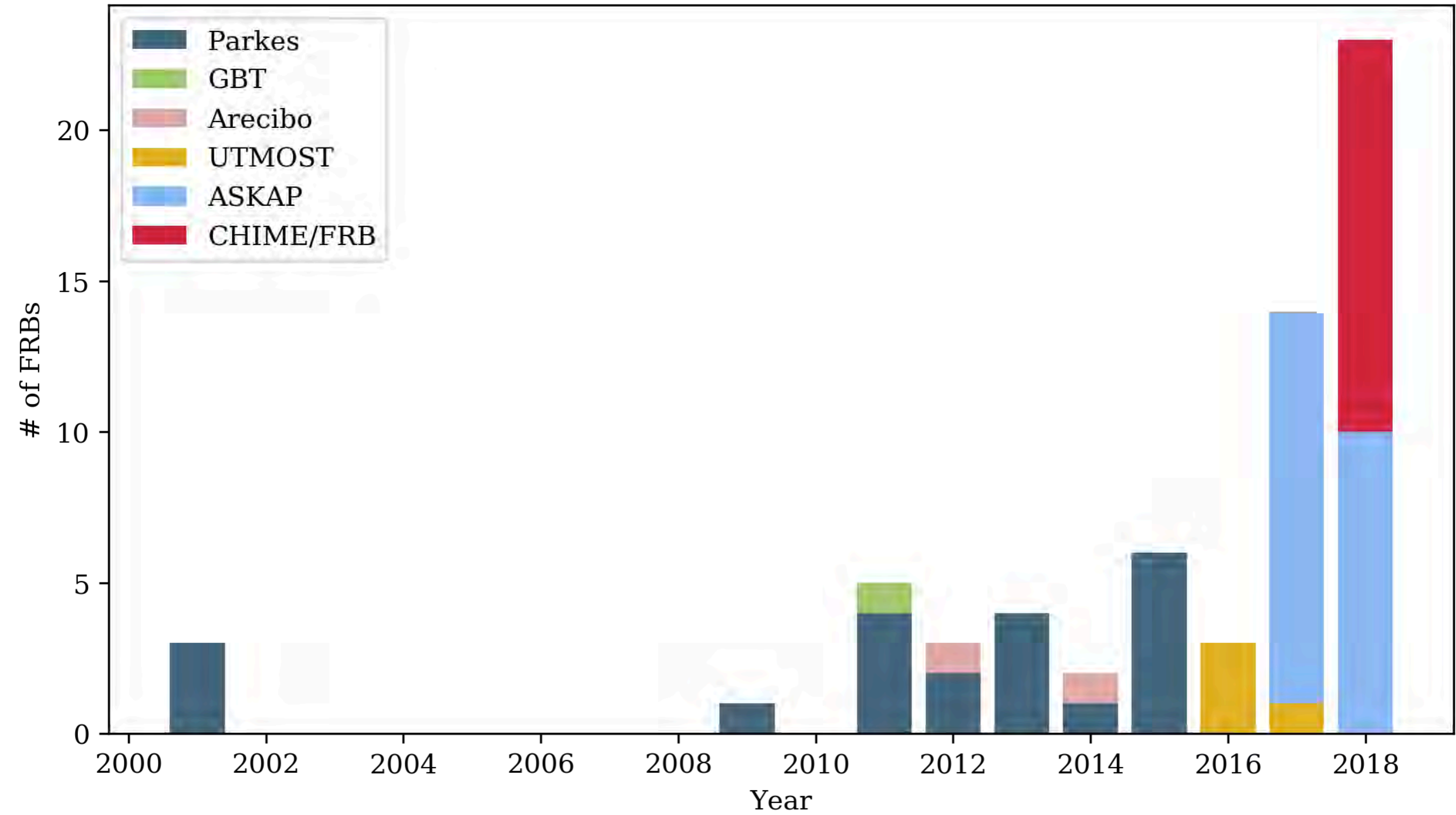


BUT, WAIT & NOTE

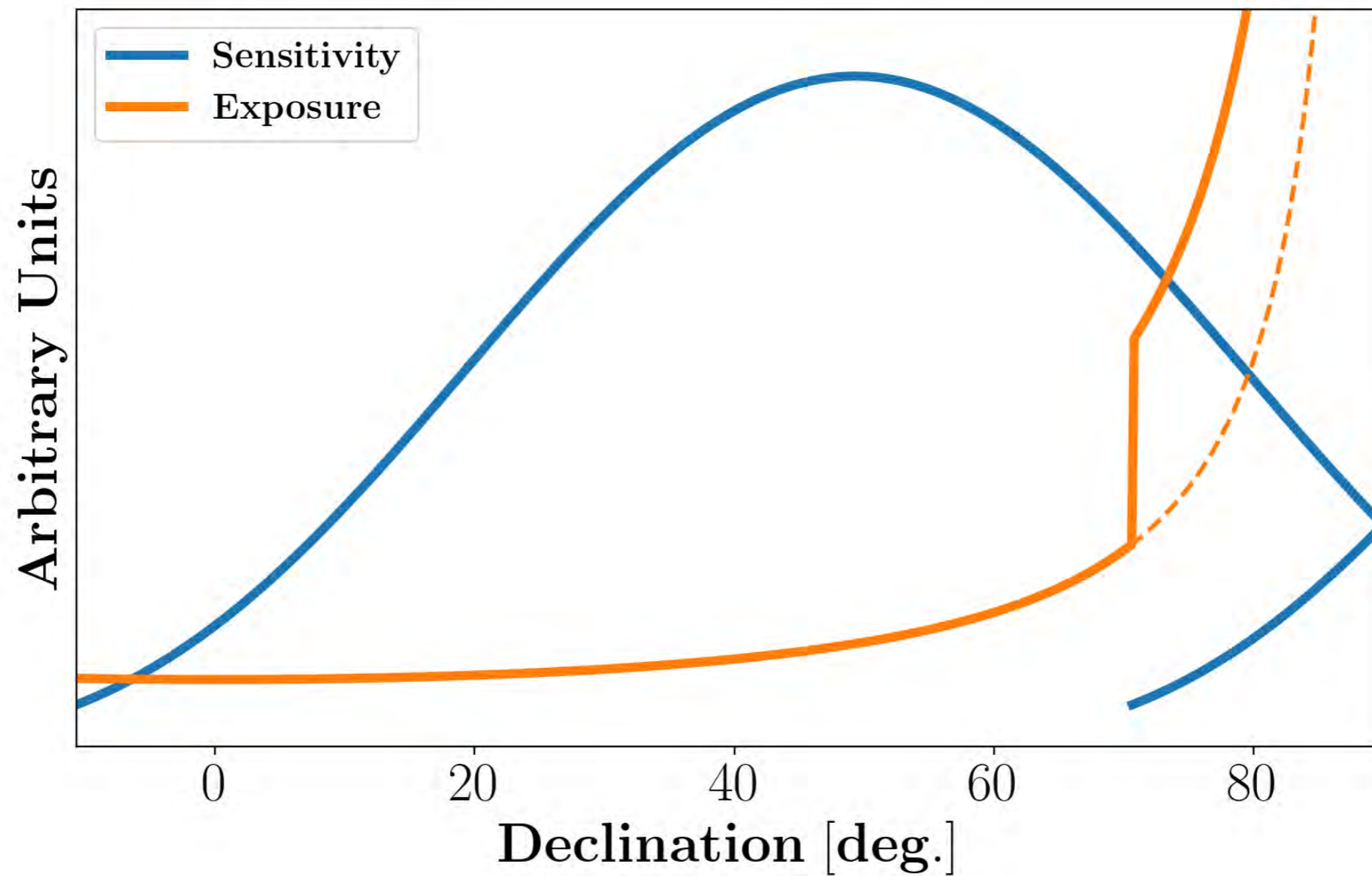
- ▶ Based on ~1 month of engineering data, at subscale capacity.
- ▶ No polarization data yet, uncalibrated, can't measure absolute frequency spectrum.
- ▶ Angular resolution currently suboptimal
- ▶ Selection function not well characterized
- ▶ Instrument is still in commissioning!







SENSITIVITY AND EXPOSURE



Note: sensitivity curve is just a toy model!

DM DISTRIBUTION

WHATS NEXT FOR CHIME/FRB?

- ▶ Finish commissioning & Start FRB rate experiment
- ▶ Study 250+ FRBs
 - DM, Scattering, DM Index & Sky distribution
 - Pulse Morphology
 - Multi-wavelength counterparts
- ▶ 20+ new radio pulsars and RRATs!
- ▶ Commensal Slow Pulsar Search
- ▶ Outrigger Telescopes