

GLASS and other deep fields with uGMRT



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GAMA Legacy ATCA Southern Survey - GLASS



- Deep, wide field survey of GAMA G23 field with ATCA at 5.5GHz and 9.5GHz
- Sensitivity: $\sim 30 \mu Jy$ rms at 5.5GHz and $\sim 50 \mu Jy$ rms 9.5GHz
- 50sq degree region (centre: 23h, -32.5°)
- ~2500 hours total (700 hours this semester, to finish last region and 'patch' up pointings)

G23 has outstanding multi-wavelength coverage Deep UV, optical, nIR, far-IR (GALEX, KiDS, VIKING, WISE, HERSHEL)



GLASS Science Goals



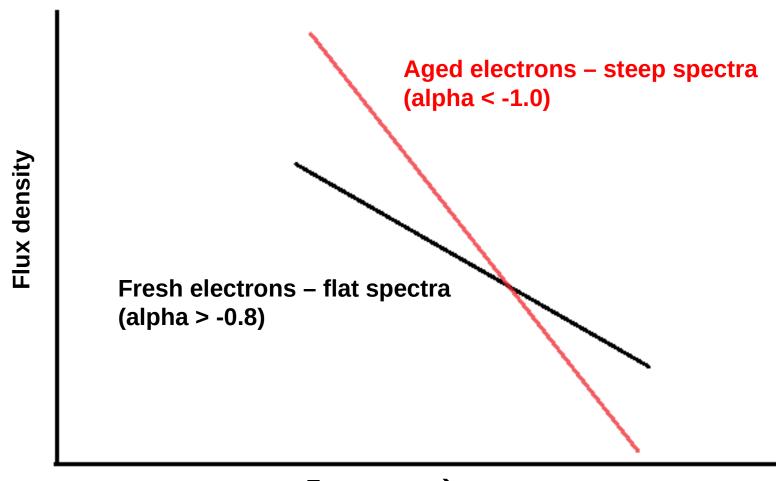
- Life cycle of radio-loud AGN
 - Double-double or 3C388 type
 - "Dead" AGNs (in Radio)



MHz vs GHz -advantage uGMRT



Steepening of synchrotron spectra with time.... stronger at low frequencies; faint at high frequencies





GLASS Science Goals



- Life cycle of radio-loud AGN
 - Double-double or 3C388 type
 - "Dead" AGNs (in Radio)
 - Atypical radio sources (X, S, Z shaped)
- Evolution of radio-loud AGN
 - -CSS/GPS sources
 - Giant Radio Galaxies
- Population of high redshift radio-loud AGN
 - -USS sources, IFRS as candidate HzRGs..
- •RLF and source count at low frequencies



GLASS TEAM



Minh Huynh (PI and GLASS Exec)	CASS and ICRAR/UWA	
Nick Seymour (GLASS Exec)	ICRAR/Curtin	
Stas Shabala (GLASS Exec)	U Tas	
Kate Chow (Project Manager)	ICRAR/UWA	
Luke Davies	ICRAR/UWA	
Martin Meyer	ICRAR/UWA	
Ray Norris	WSU/CASS	
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Isabella Prandoni	IRA/INAF	
Thomas Franzen	ASTRON	

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Gulay Gurkanuygun	CASS				
Chris Riseley	CASS				
Andrew Butler	ICRAR/UWA				
Tim Galvin	CASS				
Jordan Collier	UCT				
Andrew O'Brien	WSU, CASS				
Ross Turner	UTas				
Jesse Swan	UTas				
Josh Marvil	NRAO				
Jonathan Rogers	UTas				
Ishwara Chandra	NCRA-TIFR				
Dharam Lal	NCRA-TIFR				
Biny Sebastian	NCRA-TIFR				
Manager Day					

Veeresh Singh, Mousumi Das...



uGMRT band 3 observations (250–500 MHz) (Cycle 32, August 2017)



- 30 microJy/beam at 5.5 GHz = 190 microJy/beam at uGMRT band 3 (250 500 MHz), for a spectral index 0.7
- Requested 66 hrs to cover 53 pointing, awarded 33 hrs.
- Observations carried out using bandwidth 200 MHz (300 500 MHz), 4K channels and 8sec integration.
- about 30 mins/pointing (10mX3 scans semi-snapshot)
- Analysed using CASA-based pipeline
 (http://www.ncra.tifr.res.in/~ishwar/pipeline.html)



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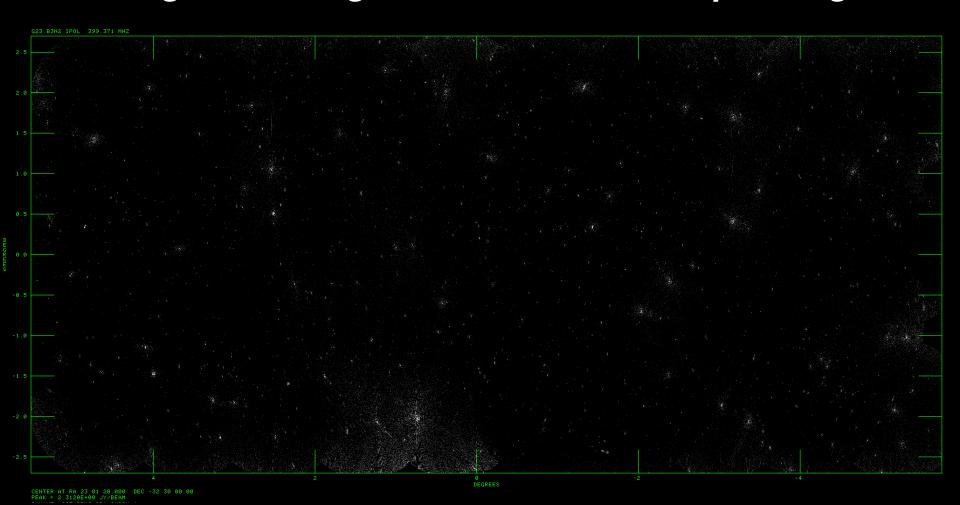




The ATCA-GAMA G23 Legacy Field



Preliminary image from uGMRT available! 12Degree X 5 degree MOSAIC from 50 pointings

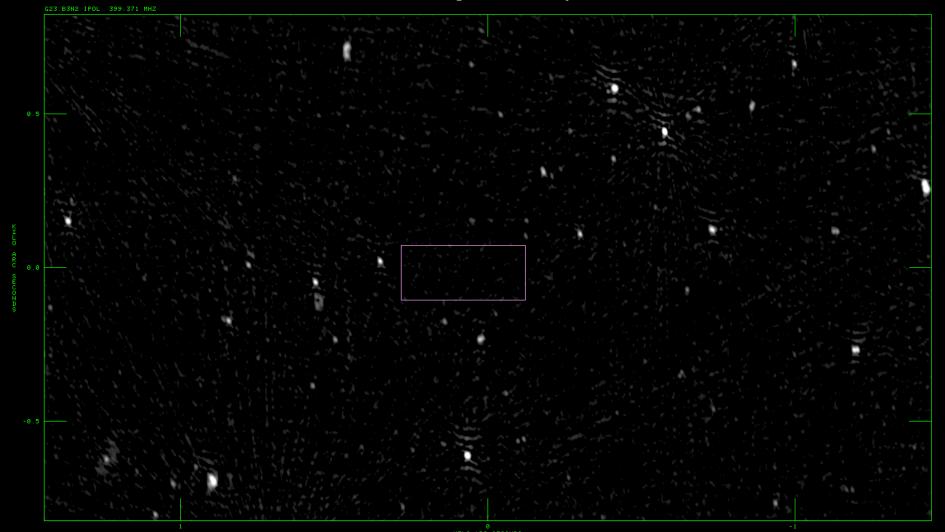




The ATCA-GAMA G23 Legacy Field



A small region.....
rms ~ 130 microJy/beam; resolution ~ 10"





uGMRT followup

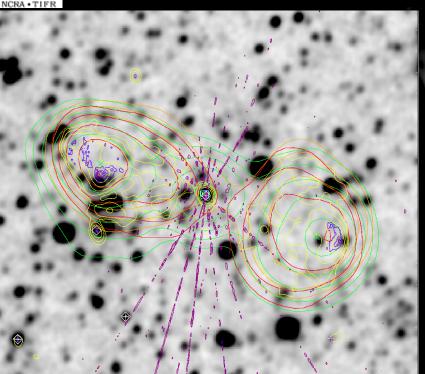


Subsequent proposals with uGMRT in 2018 and 2019 for individual sources in band-3, band-4 and band-5.



Giant Radio Source





TGSS, MWA Overay

uGMRT 400 MHz

uGMRT 680 MHz

Seymour et al, PASA, 2019.

ATCA 5.5 GHz



Possible Indo-Aus collaborations



GAMA and G23:

uGMRT band-3 (250 – 500 MHz) preliminary images of G23 field (50 sq degrees) available and improvement coming *soon*!

uGMRT band-4 (550 – 850 MHz) and band-5 (1000 – 1450 MHz) followup observations of selected sources from G23 completed with science ready images

New: Equitorial GAMA fields with uGMRT?



Possible Indo-Aus collaborations



uGMRT neatly fills the gap between MWA and ASKAP! (high freq for MWA and low freq for ASKAP!)

Low freq counterpart to EMU suvery at uGMRT band-3

- Matched resolution and sensitivity
- Pilot proposal in the next GTAC Cycle planned

Deeper TGSS in uGMRT band-2 over GLEAM-X?

Looking forward to more discussion