



Image Courtesy: Quasar ULAS J1120+0641, ESO/  
M.Kornmesser

ISTRAC/ISRO

# Delta-DOR Correlator Updates & Future Plans

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# AGENDA

- Overview
- Correlator
- Recent Updates
- Future Plans

## Delta Differential One-Way Ranging (DDOR)

- VLBI-based Spacecraft Navigation Technique
  - Pioneered by NASA/JPL in the 1970s and 80s.
- Evolved for interplanetary missions
- Gets accurate angular position of a Spacecraft in the nano-radian range
  
- Supplements the Ranging and Doppler Navigation Techniques
  
- Crucial especially after Orbit Insertion Maneuvers as with ISRO's Mars Orbiter Mission (MOM)
  - Typically - brings accuracy of orbit-determination in the 1-2 kms range at Mars distance.

# Typical Mars Orbital B-Plane Error Ellipses

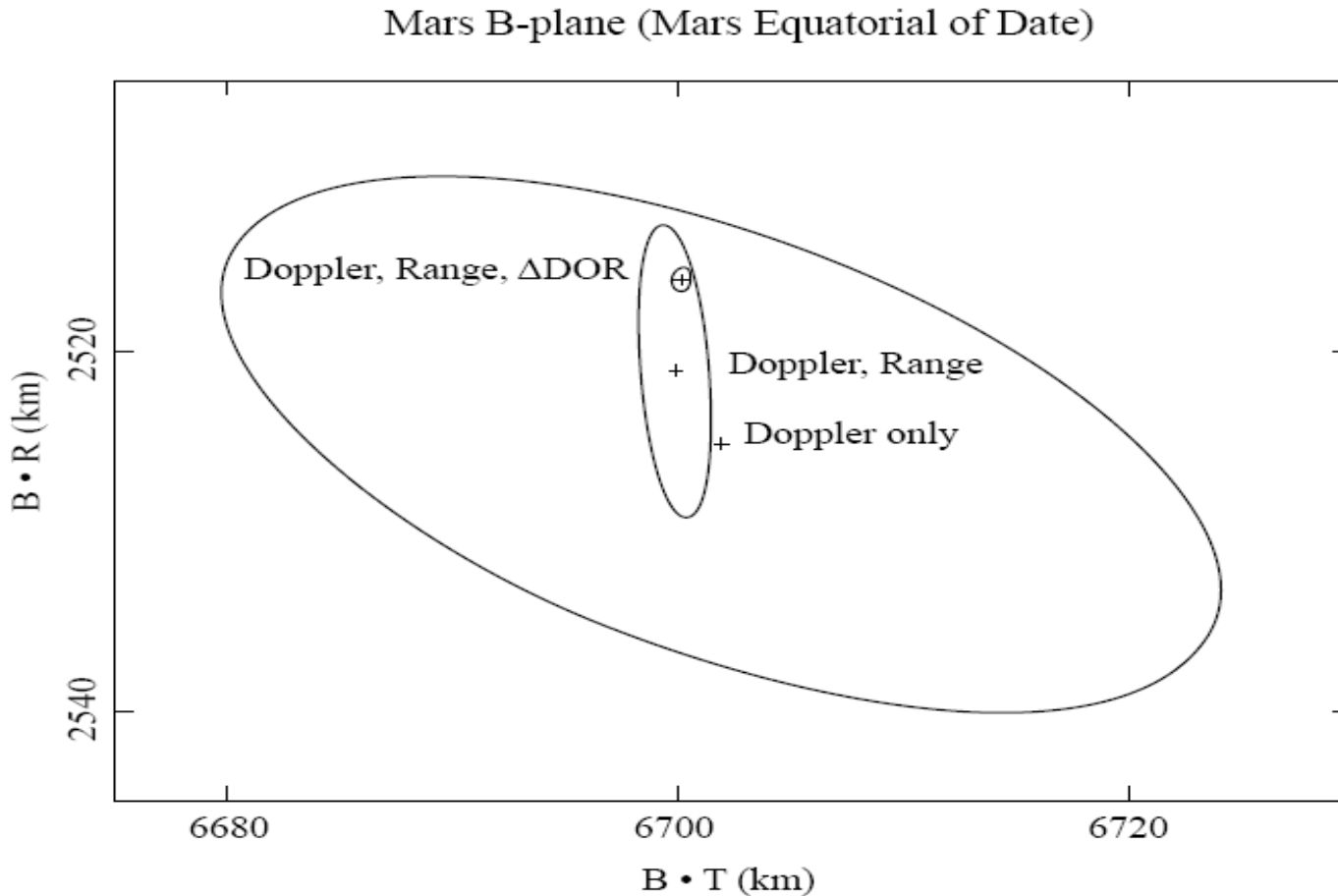


Figure 2-2: Error Ellipses in the Mars Targeting Plane<sup>1</sup>

<sup>1</sup>Courtesy: JPL/CalTech

# $\Delta$ DOR Concept

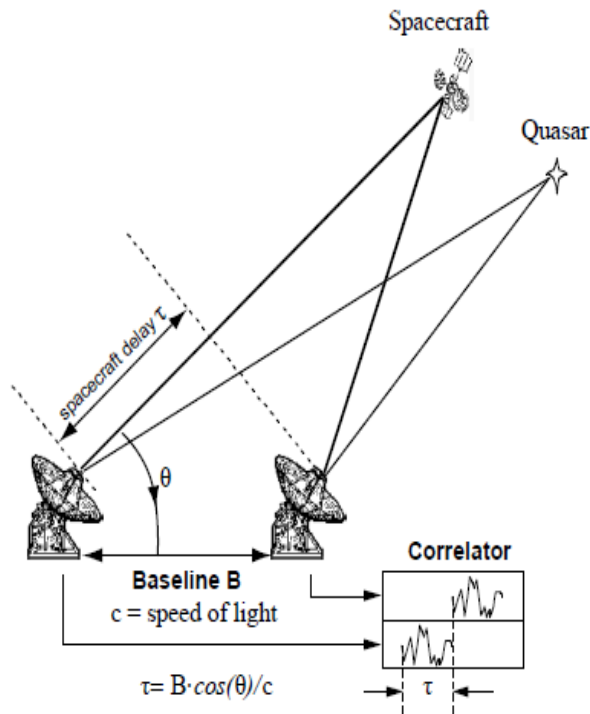


Figure 2-1: Delta-DOR Observation Geometry

Figure Courtesy: CCSDS RDEF Blue-Book

- $\tau_{sc}$  Differential One-way Ranging (DOR) Measurement
  - Itself tells the Spacecraft-Baseline angle  $\theta$
  - But has errors due to Station Clock Errors and other Instrumental and Atmospheric media effects
- $\tau_{qsr}$  DOR Measurement of well-known catalogued radio-sources : Quasars, are used to derive and eliminate these common errors and improve accuracy of  $\theta$
- QSR: Essentially, a Calibration Source for the Measurement

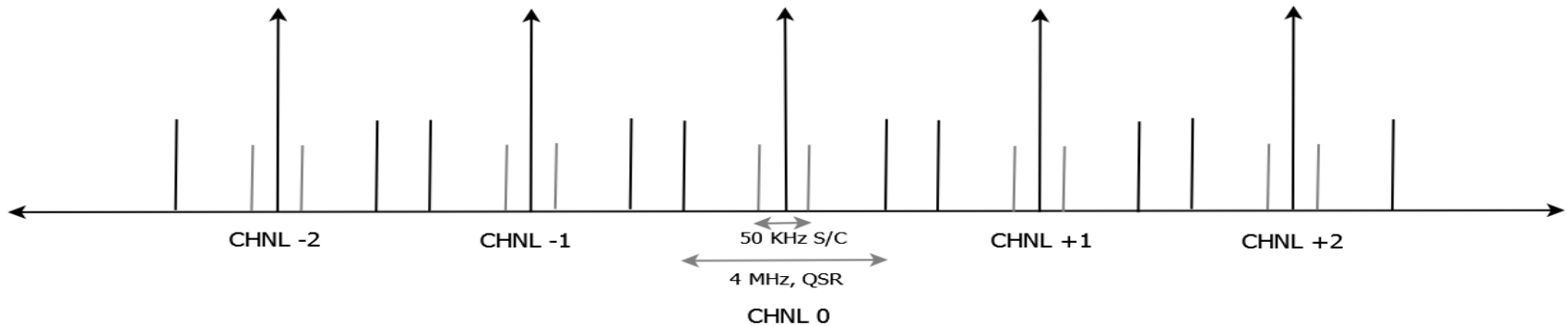
# Diff. One-Way Delay: Qsr vs S/C

- **Quasar data** is wide-band Gaussian noise, and is well below the station system noise-floor
  - Correlation and integration of the data for large number of samples are required to extract out the Quasar data
  - **S/C signal** is well above the station system-noise floor and is visible during recording
  - It is also a well-defined sinusoid
  - Need is to establish a PLL type of mechanism to extract the phases at both the stations and get the differential phase
- Hence the optimal algorithms to derive the Differential One-Way delay for S/C and QSR are different.

# Typical DDOR Session

- DDOR Sessions are planned in a sequence as follows:  
S – Q1 – S – Q2 – S – Q1 – S – Q2 – S
- A session lasts about an hour, and each scan S/C or QSR around 5 to 10 mins
- This sequence allows interpolation of measurements to the required time of S/C or QSR scan
- Q1, Q2 are generally chosen to within 10 degrees of the Spacecraft.

# DDOR Channel Recording



- Central channel set to 2296.3 MHz + Predicted Doppler at both Stations
- Harmonics are 3.9 MHz apart [There actually is slight overlap between quasar channels: not shown in picture for clarity]
- 50 KHz Bandwidth, 8-bit complex I&Q sampling for S/C
- 4 MHz Bandwidth, 2-bit complex I&Q sampling for QSR
- Data recorded as per the CCSDS Raw Data Exchange Format (RDEF)
- Open specification – standard format for multi-agency DDOR data.



# ISRO's Mars Orbiter Mission (MOM) ..

- ◉ MOM carries a DDOR transmitter module
- ◉ DDOR sessions of MOM NASA/DSN stations at
  - ◉ Canberra (CNB)
  - ◉ Madrid (MAD) and
  - ◉ Goldstone, California (GDS) where scheduled
- ◉ DDOR-results were crucial in ensuring successful Mars Orbit Insertion on 24<sup>th</sup> Sept 2014
- ◉ Later, DDOR recording capability added to the Indian Deep-Space Network Station (32m) dish at Byalalu, near Bangalore
- ◉ BLR-MAD and BLR-CNB sessions have been successful thereafter..

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# Delta-DOR Correlator

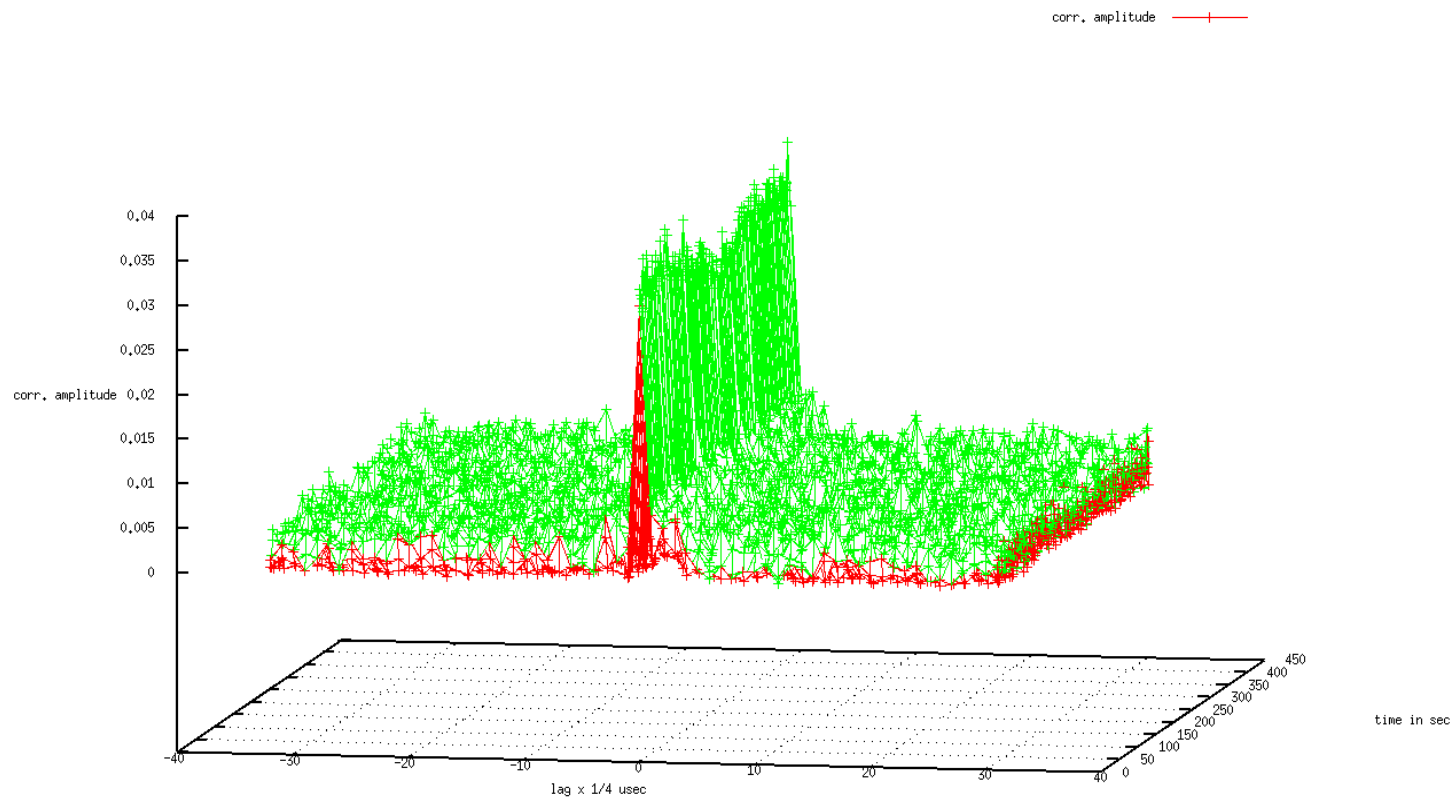
- ◉ Effort to indigenously develop our own DDOR Software Correlator.
- ◉ Entailed Three Major Aspects:
  - ◉ Quasar Correlator
  - ◉ Spacecraft Correlator
  - ◉ Model-Delay Generator

# Quasar Correlator

- XF-model
- First-level cross-correlator
- Implements Fringe-Fitting via Bandwidth Synthesis – using all the 5 channels
  - via both
    - MIT HOPS Fourfit and
    - JPL PhaseTracking Approaches

# Sample First-level Correlation Results

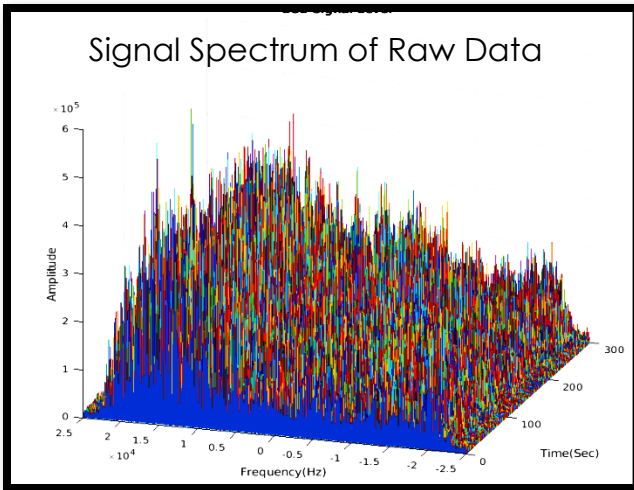
2015 D0Y: 053, P\_2355-106, Chnl-0 corr-amplitude



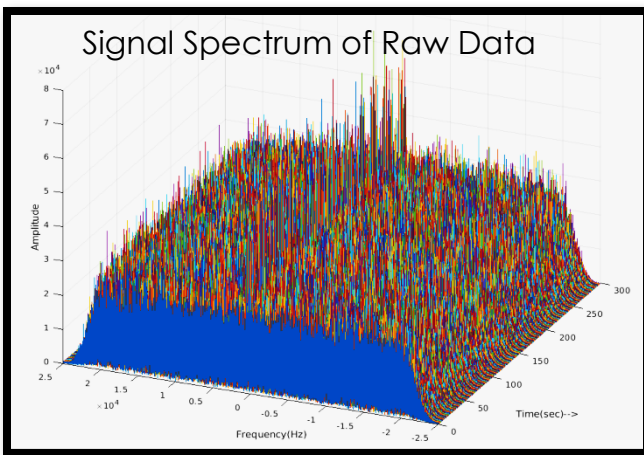
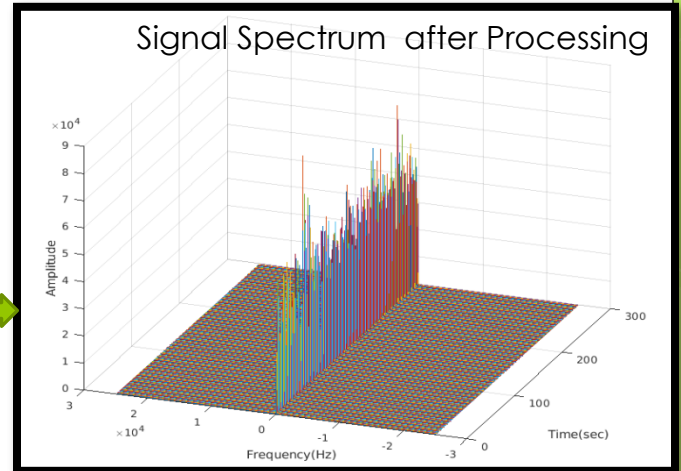
# S/C Correlation

- General PLL based approaches such as Analog / Digital PLL, AGC-based PLL approach, Complex-Filtering, Windowing-based Digital Filtering method etc., alone were not sufficient to handle the OUTER channels (+2, -2 harmonics) noisy-data processing.
- Evolved our own approach to doing accurate phase measurement in this environment

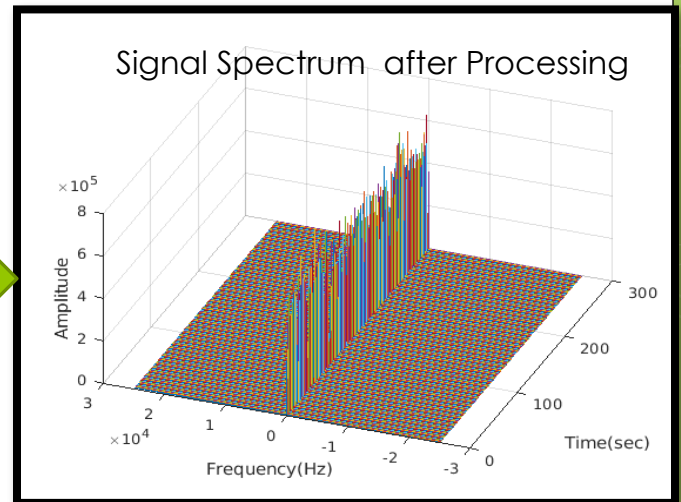
# Outermost S/C Channel Processing



Exact Signal Extraction  
From D32 +2 Channel  
Extraction



Exact Signal Extraction  
From D32 -2 Channel  
Extraction

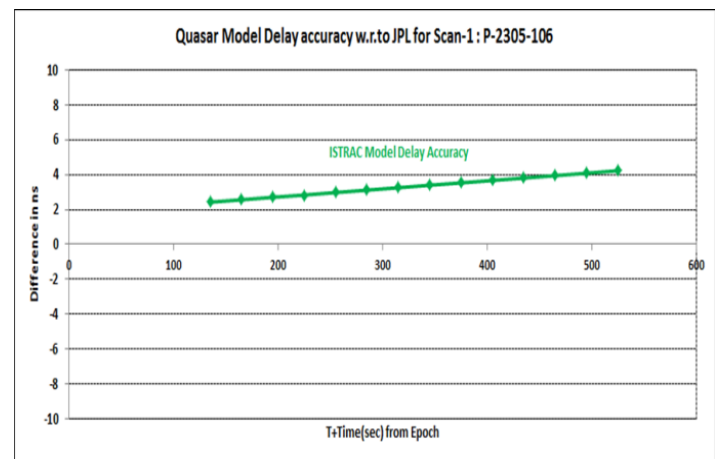
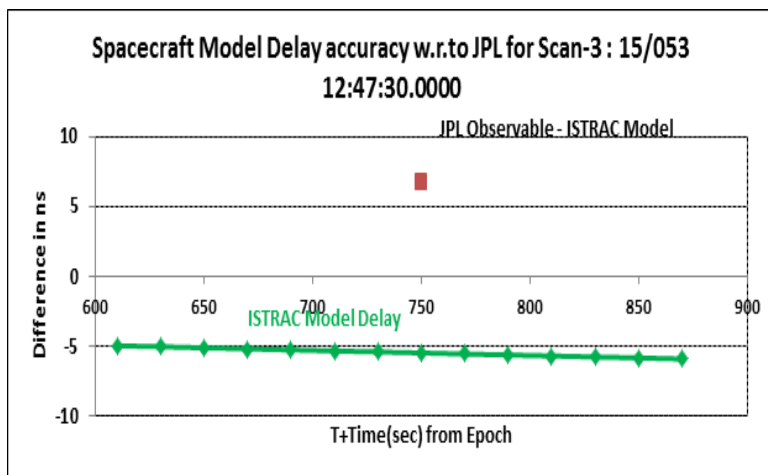
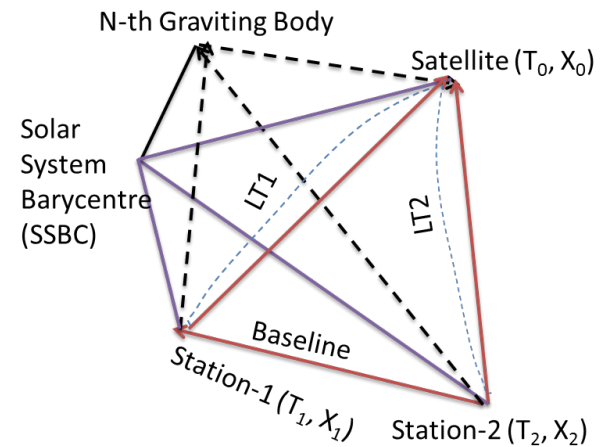
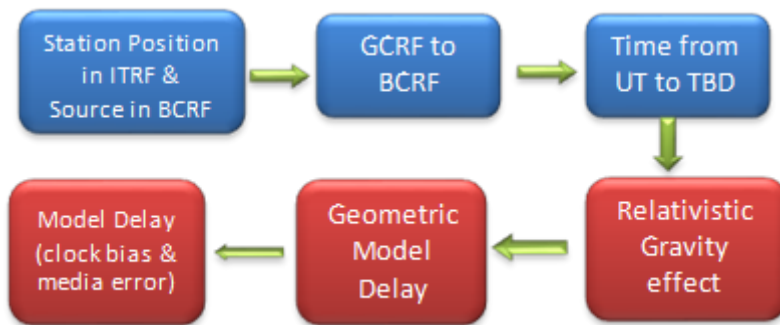


# Delay Model Generation

- Accurate Delay Model Generation – on the order of 10-20 nsec is mandatory for accurate Quasar Correlation
- Initially adapted DfxCALC-11 for QSR purpose
- S/C Correlation also requires good modeling of S/C expected delay.
- CLOCK-BIAS estimation based on Quasar Correlation is an essential step in this process - also getting completed.
- Independent ISTRAC delay-model is developed for both QSR and S/C and yielding good results.



# Delay Model Generation



# Correlator Status (Nov 2019)

- Getting good observables
  - Matching at sub nano-sec level with those produced by NASA/JPL Correlator for both Quasar and S/C DOR.
  - Delay-Modeling is good and meeting requirements.
- Correlator overall is satisfactory for Operations.

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# Clock-Bias (cbias) Estimation

- Good Frequency Standards: H-MASER, Caesium
  - But **micro-sec** order clock-offset with UTC (therefore, also with each-other) is typical
- Critical to estimate this for successful correlation

## Idea

- Quasars are stationary for all practical purposes;
  - Locations very well known (sub milli-arc-sec level)
  - Use the strength of Quasar Delay-Modeling (accurate in **10-nsec** order) to estimate cbias
- 
- Clock-Bias search is the initial step to correlation
  - Delay models (QSR & S/C) are updated with Clock-Bias before correlation

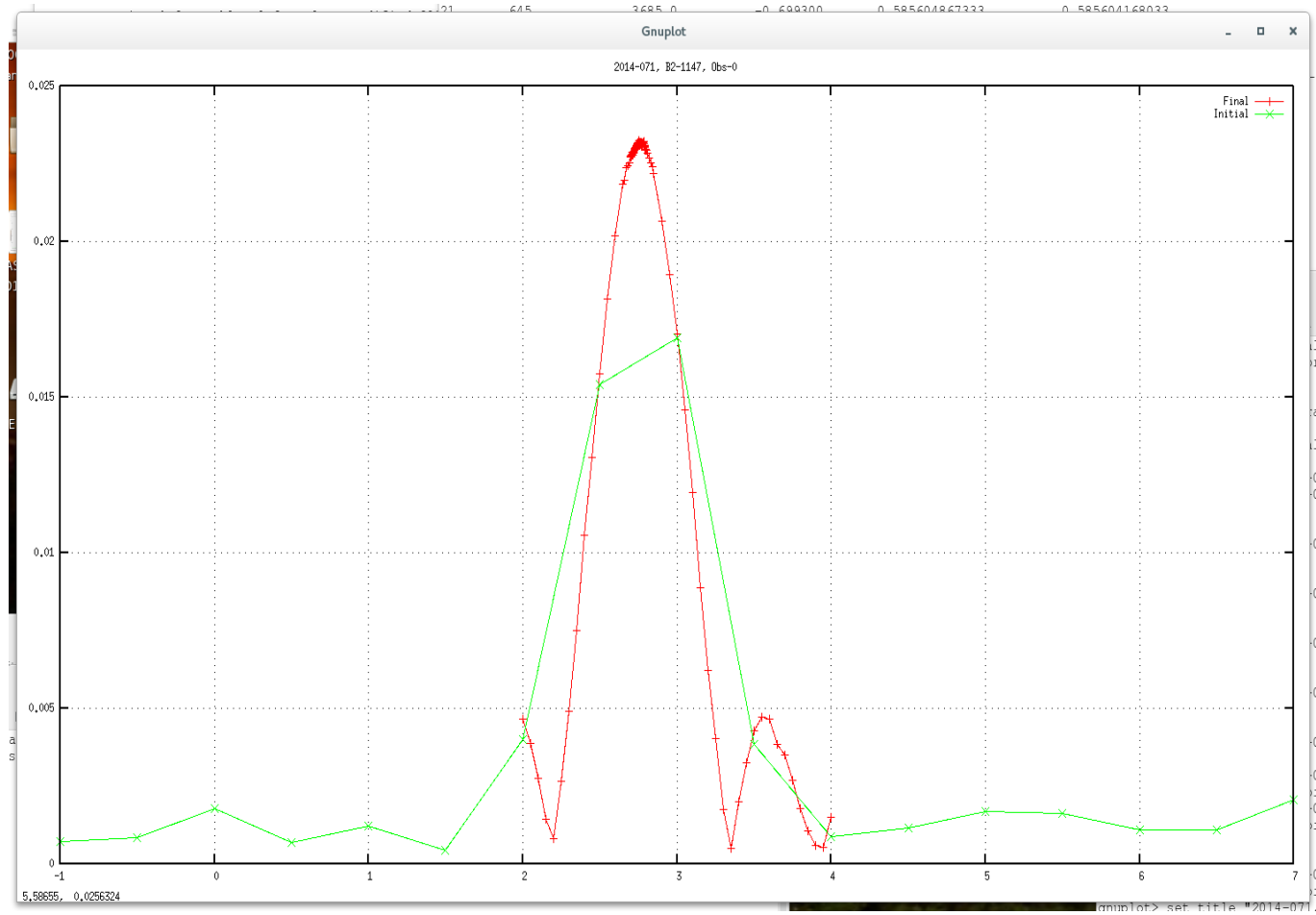
# CBIAS Estimation

- Quasar Sampling is done at 4 MHz: this is 0.25 usec or 250 nsec interval
- We need an estimate accurate in 10 nsec order
- Initial Quasar Correlation:
  - Choose lag-window large enough to cover max. CBIAS expected – may be  $\pm 5$  usec
  - A shorter period – 30 sec – of data, should suffice
- Perform Correlation and locate the lag value where correlation peak-amplitude is seen. This is the initial guess for the iterative procedure.

# CBIAS Estimation (contd.)

- Iterative Procedure:
  - Try trial cbias values – say  $\pm 5 \times 100$  nsec intervals around the initial-guess value for the lag. Locate the peak in this iteration.
  - Try trial cbias values – say  $\pm 5 \times 10$  nsec intervals around this peak. Locate the peak in this iteration
  - Try trial cbias values – say  $\pm 5 \times 1$  nsec intervals around this peak. Locate the peak in this iteration
  - Fit all the correlation amplitudes vs lag trial-values with  $\text{abs}(\text{sinc}())$  function, and estimate peak.
  - Average the peak values obtained for all channels to give final estimate
- This is our best-estimate of cbias and is typically accurate to 10 ns order

# CBIAS Results (Sample)



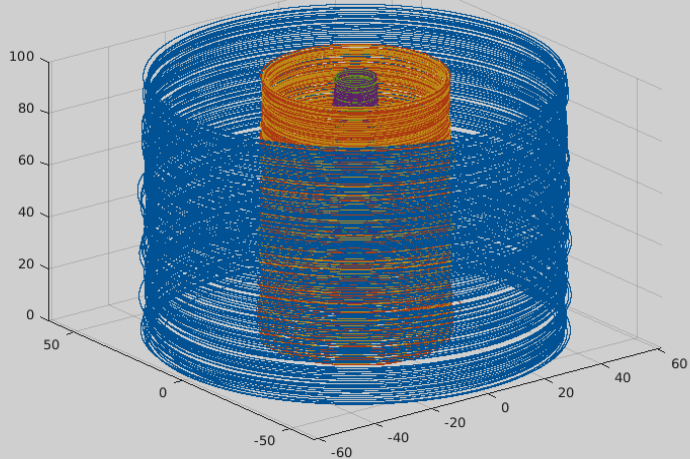
# S/C Correlation

- ◉ Initial validation done for DDOR Sessions in orbital-phase
- ◉ Now, validated our Correlation for DDOR Data for the more critical cruise, pre-Orbit-Insertion phase
- ◉ Here, continuous tuning of local-oscillator wrt predicted-Doppler frequency was done for higher accuracy
- ◉ Spacecraft correlator software was modified for this.
  - ◉ It is able to produce consistent results in this scenario as well.

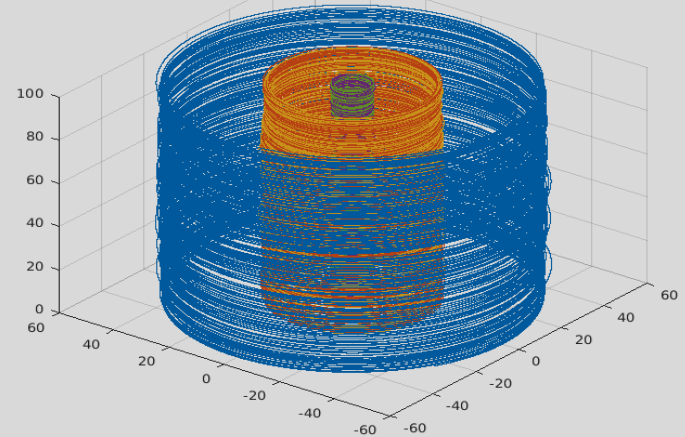


# Results for Just Before MOM MOI Data Sets

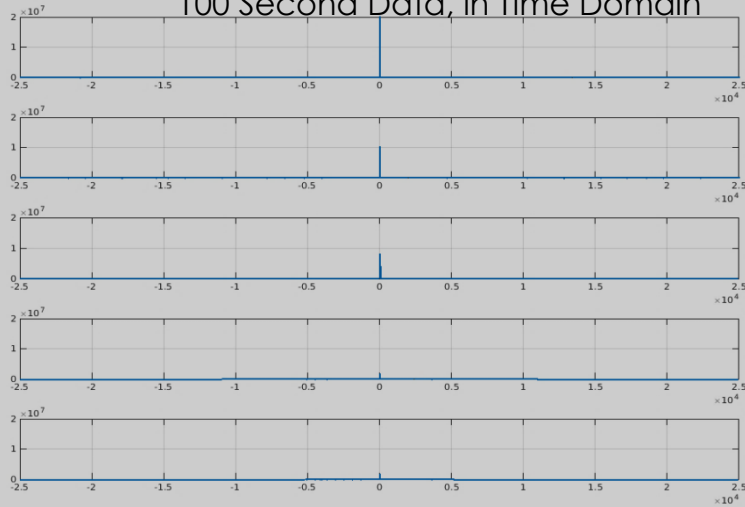
D15 All 5 channel's Complex Signal  
(After Processing)



D45 All 5 Channel's Complex Signal  
(After Processing)

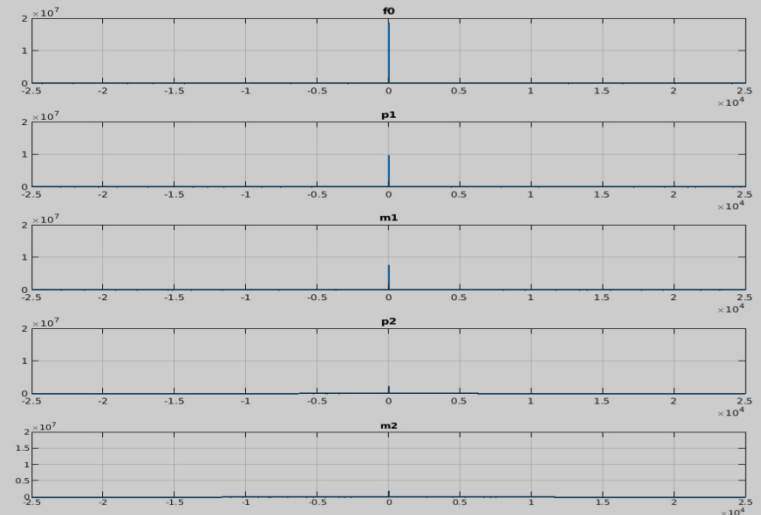


100 Second Data, in Time Domain



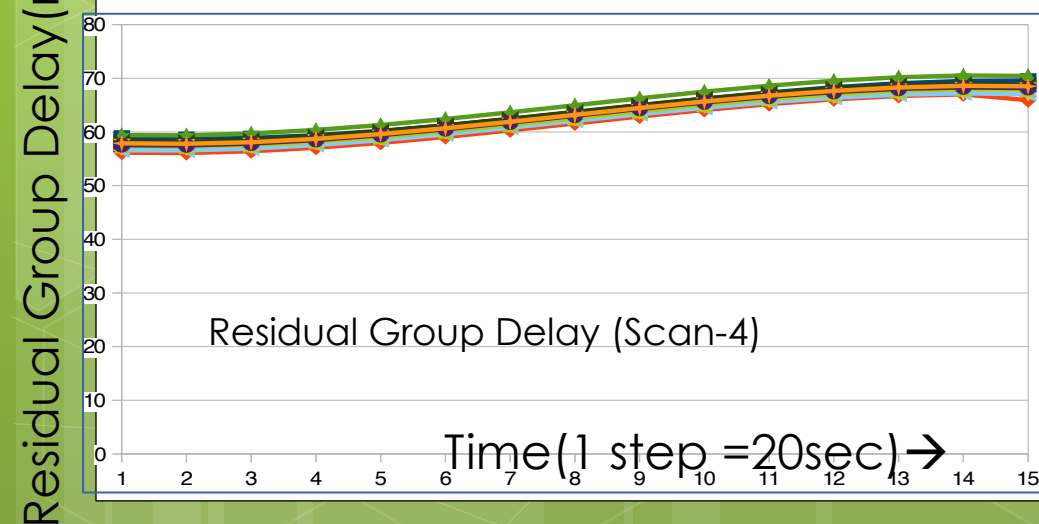
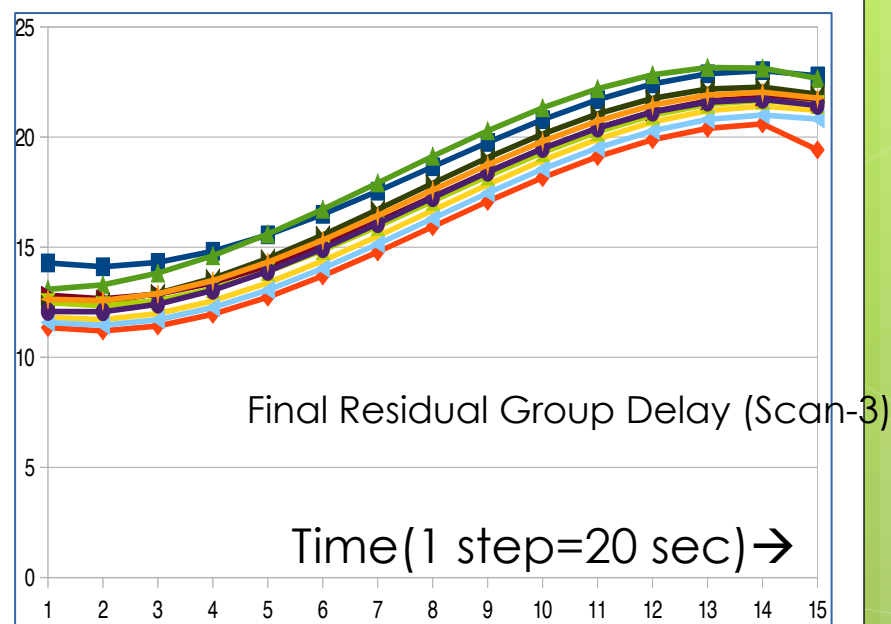
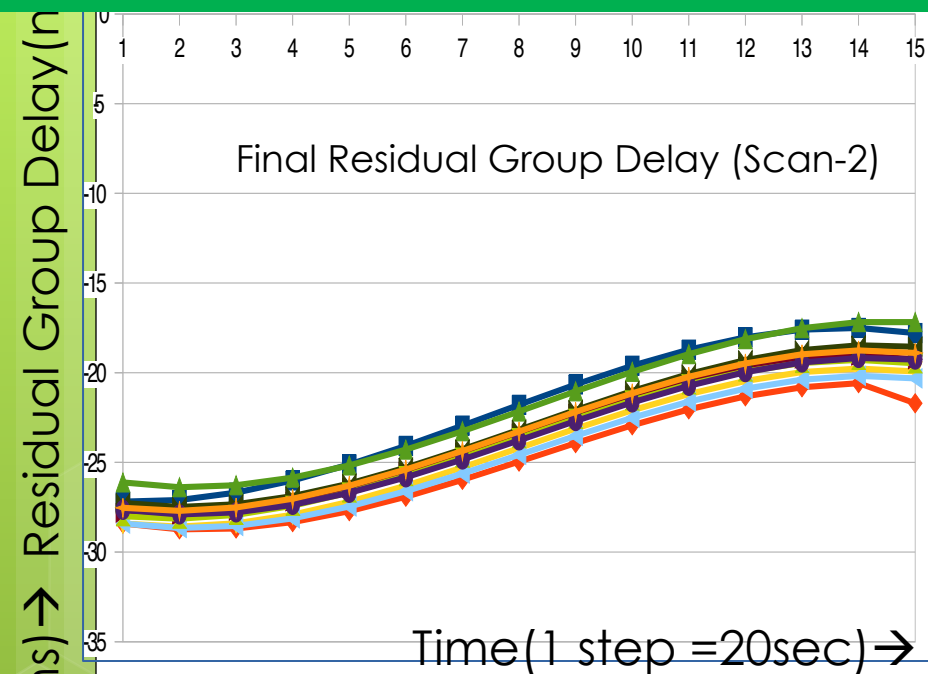
Spectrum of all five channels of D15

100 Second Data, in Time Domain



Spectrum of all five channels, D45

# Results for Just Before MOM MOI Data Set



Remarks :  
Residual Phase Delay  
Computation using all channel's  
data for all the three spacecraft  
scans shows good results.

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# Performance Optimization

- Correlation is time-consuming!
  - Optimize for a multi-core / GPU operational environment: **Need for Speed**
    - **3 to 3 ½ hours'** processing may be brought to **~ ½ hour**
- Basic parallelization: Correlate Each Channel independently as a thread
- Clock-Bias Search – each clock-bias trial value can be run independently as a thread
- Complex multiplication over a lag-window to be done as Single-Instruction-Multiple-Data (SIMD) Instruction
- Each second's correlation can be done independent of the next second's correlation – as much as CPU / Memory resources support..

# Future Plans Contd..

- Using VLBI for other *solve-for* parameters such as Station Location estimation.
- Format conversion – CCSDS RDEF  $\leftrightarrow$  VLBI VDIF format and sharing of data from joint observations..
- IDSN-32m station – S/X Band participation in VLBI experiments – subject to the operational schedule and management clearances..
- Media Calibration – Best Practices / Approaches